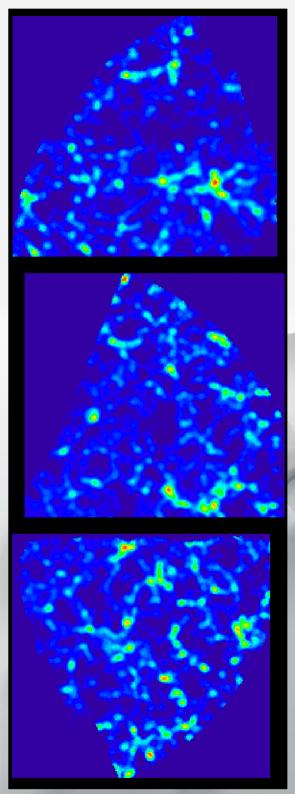
Future Virialized Structures

An analysis of superstructures in SDSS-DR7



Largest gravitationally bound systems (presently non-virialized) Capricornus Corona-Borealis Supercluster Complex morphology: very luminous and colors cores, connected by Supercluster filaments Hercules **Superclusters** Capricornus Void Corona Borealis Lengths of 50-150 Mpc Boötes Void **Superclusters** Microscopium Pavo-Indus Void Supercluster Boötes Void Centaurus Shapley Supercluster Supercluster Hydra-Centaurus Sculptor Void Supercluster Sculptor **Superclusters** Virgo Superclus Pisces-Cetus **Superclusters** Hydra Coma Supercluster Ursa Major Perseus-Pisces Supercluster Supercluster Phoenix Fornax Supercluster Leo Void Canes-Major **Superclusters** Void Columbia Void Sextans Horologium Supercluster Columba Supercluster Supercluster



Identification Process

FoF algorithm on galaxy clusters (Einasto et al., 1997)

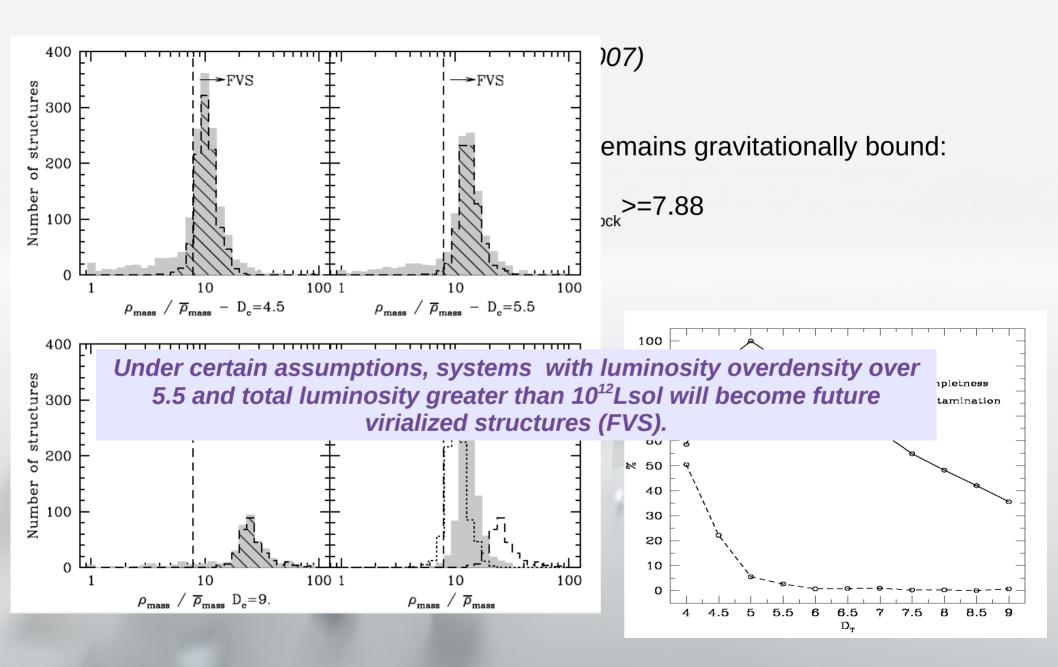
Percolation parameter?

Luminosity-Density Maps (Einasto et al., 2007)

- 1. smooth luminosity of galaxies --> Kernel
- 2. estimate luminosity density --> grid
- 3. connect overdense cells

Luminosity overdensity threshold?

Threshold selection



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Sample	Z _{max}	$D_{max}[h^{-1} Mpc]$	M_r^{llm}	Volume $[10^7 (h^{-1} Mpc)^3]$	N_{gal}	$\bar{\rho}_{him}[10^8L_{\odot}/Mpc^3]$	F	Corrected $\bar{\rho}_{hm}[10^8 L_{\odot}/Mpc^3]$
S1	0.10	293.92	-20.05	1.85	94271	0.80	2,11	1.68
S2	0.12	351.34	-20.47	3.17	89513	0.58	2.98	1.73
S3	0.15	436.55	-21.00	6.01	62344	0.29	5.66	1.64
S2c	0.10	293.92	-20.47	1.85	51188	0.56	2.98	1.73
S3c	0.10	293.92	-21.00	1.85	17507	0.27	5.66	1.64
M_{Rsp}	0.12	351.34	-20.47	3.17	106604	0.75 0.75	2.98 2.98	2.23 2.23
_		AT Z		E -l		0.75	2,70	

Sample	N_{FVS}	F_{vol}	F_{lum}	$glxs_{inFVS}$
S1	67	1.08%	10.85%	9707
S2	150	1.26%	13.54%	11394
S3	412	1.66%	20.61%	11682
M_{Rsp}	227	1.62%	18.87%	19265
M_{Zsp}	181	1.35%	15.14%	15368

Table 2. Main results obtained for the samples of identified FVS. For each sample, we show the number of future virialized structures N_{FVS} , the percentage of volume occupied by FVS F_{vol} , the percentage of luminosity of galaxies within FVS F_{lum} and the total number of galaxies within FVS $glxs_{lnFVS}$.

de in the r-band is in the range $14.5 \le r \le 17.5$. The mean ixies. The correction factor F (Eq. 6) and the resulting mean



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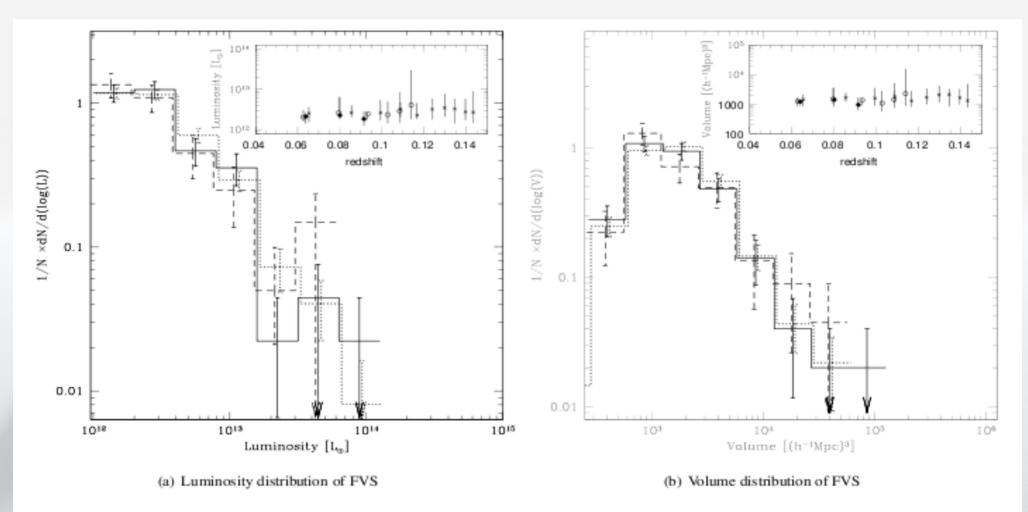


Figure 5. Luminosity and volume distributions of structures for samples S1 (dashed line), S2 (solid line) and S3 (dotted line). The insets show the average volume and luminosity as a function of the redshift, in redshift bins with equal numbers of superstructures, with filled circles, empty circles and cruxes, respectively for samples S1, S2 and S3. Error bars on the histograms indicate Poisson uncertainty. In the inset box, the uncertainty bars correspond to the 25 per cent and 75 per cent percentiles.

Future Virialized Structures on SDSS-DR7

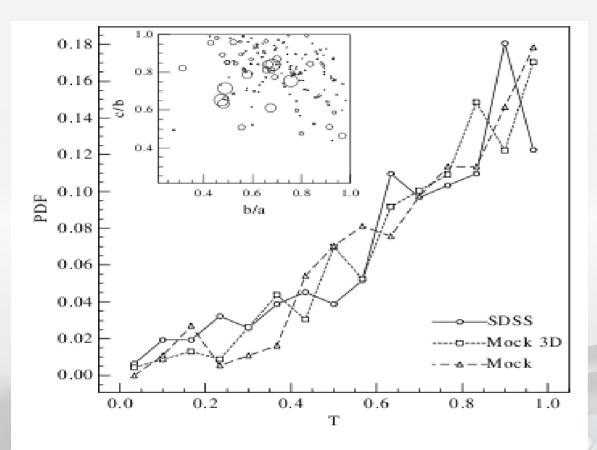


Figure 6. Probability density distribution estimates of the shape indicator parameter T (defined in section Section 4.1). The distributions correspond to FVS from the SDSS galaxy catalogue (solid line), the real–space mock catalogue (short dashed line) and the redshift–space mock catalogue (long dashed line). The inset shows a scatter-plot of the semi–axis ratios c/b and b/a that characterize the shapes of the FVS. The sizes of circles are proportional to the number of galaxies contained in each FVS.

$$T=(1-(b/a)^2)/(1-(c/a)^2)$$

T --> 1 indicates prolate structures

T --> 0 indicates oblate structures



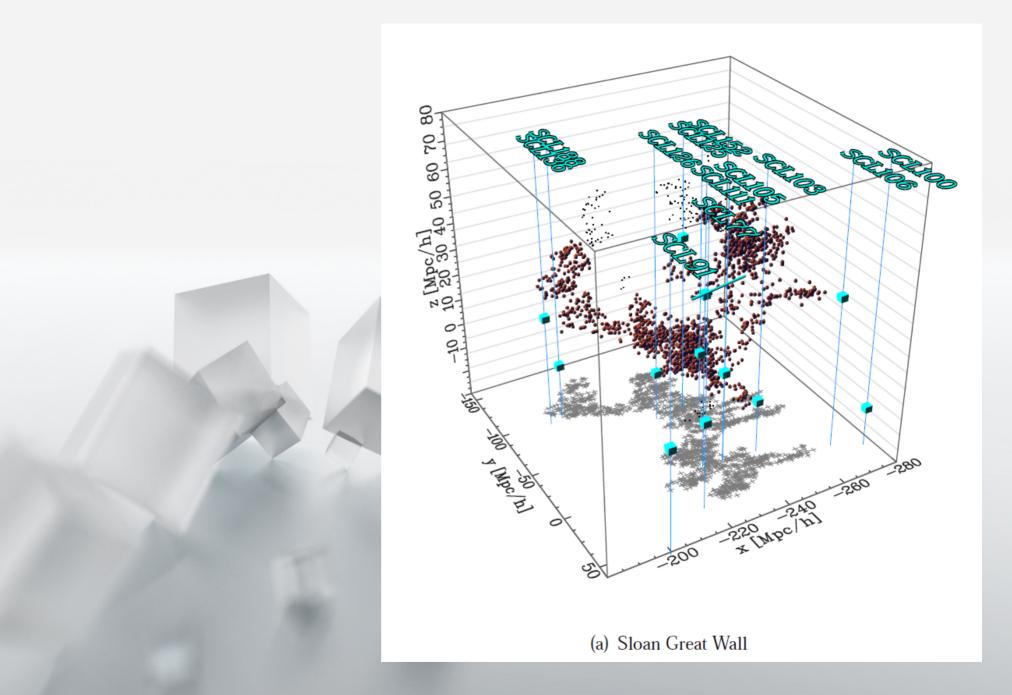
Redshift space vs. Real Space FVS

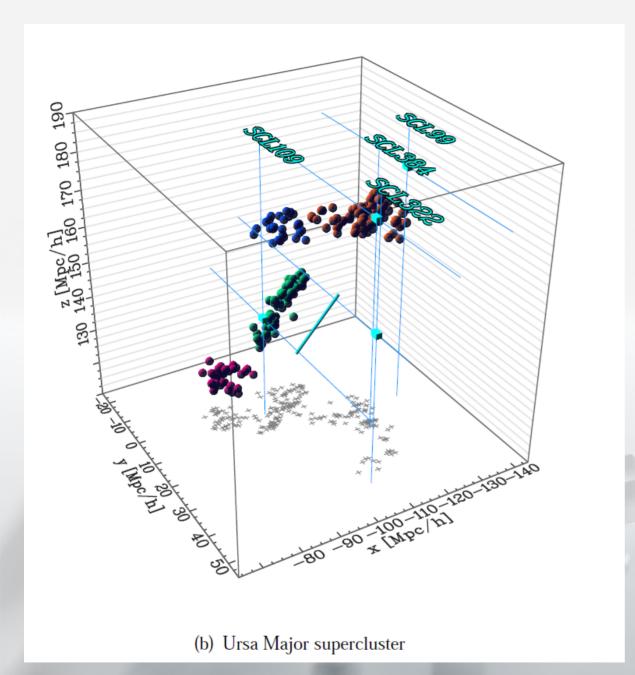
Lower number of FVS in redshift-space in comparison to real-space FVS.

20% of real-space structures are lost when analyzing redshift-space data.

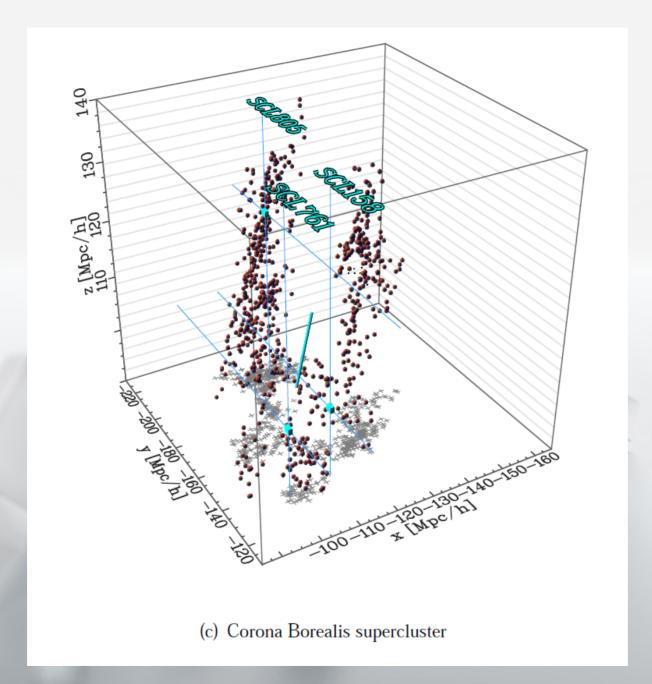
Only 2% of FVS identified in redshift-space are not real-space FVS.

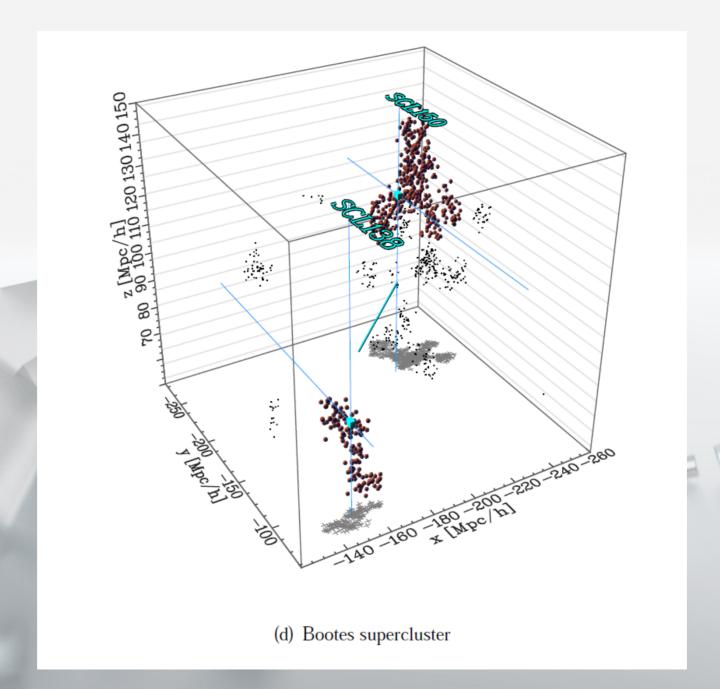
3% of redshift-space FVS are associated with more than one FVS in real-space.











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